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Plasma edge modelling with ICRF coupling

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Abstract. The physics of Radio-Frequency (RF) wave heating in the Ion Cyclotron Range of Frequencies (ICRF) in the core plasmas of fusion devices are relatively well understood while those in the Scrape-Off Layer (SOL) remain still unresolved. This paper is dedicated to study the ICRF interactions with the plasma edge, mainly from the theoretical and numerical point of view, in particular with the 3D edge plasma fluid and neutral transport code EMC3-EIRENE and various wave codes. Here emphasis is given to the improvement of ICRF coupling with local gas puffing and to the ICRF induced density convection in the SOL.

1 Introduction

The coupling of Radio-Frequency power in the Ion Cyclotron Range of Frequencies (ICRF) to the plasma depends sensitively on the plasma density in front of the antennas because the launched fast wave is evanescent below the cut-off density, usually of the order of 10^{18} m^{-3} for typical frequencies and antenna spectrum in nowadays tokamak. Recent experiments on AUG [1] and JET [2, 3] indicate that by injecting the fueling gas from the main chamber instead of the divertor, the local density in front of the antennas and thus the ICRF power coupling can be greatly increased. The corresponding EMC3-EIRENE simulations [4-6] successfully reproduce these experiments and deepen our understanding on the associated physics. This paper will review these results, summarize the mechanisms and make extrapolations to ITER and DEMO.

While generating the fast wave ($E_{\perp} \gg E_{\parallel}$, parallel direction along \vec{B}) to heat the core plasmas, the slow wave ($E_{\parallel} \gg E_{\perp}$) is also excited parasitically. The parallel electric field of the slow wave, especially close to the limiters of the antennas, is expected to be the source of the nonlinearly enhanced sheath potential in front of the antennas by accelerating the more mobile electrons out of the plasma. The inhomogeneous sheath potential drives $E \times B$ drifts in the Scrape-Off Layer (SOL) and modifies the plasma density locally. To understand this ICRF induced SOL density convection, two numerical methods have been developed: (a) simulations based on experimental data; (b) self-consistent simulations with

several codes including EMC3-EIRENE [7] (extended to include prescribed drifts [8] and plasma convection), RPLICASOL [9] (ICRF fields) and SSWICH [10] (rectified sheath potentials). This paper will review these two numerical methods and the corresponding results.

2 ICRF coupling

The local density in front of the ICRF antennas can be tailored by changing the deuterium (D2) gas puff locations. While changing the gas puff locations, no clear change of the energy confinement time and no plasma degradation are seen [3]. To understand the effects of localized gas sources on the SOL density, comprehensive 3D simulations with the EMC3-EIRENE code on AUG [4], JET [5] and ITER-relevant [6] scenarios have been carried out.

In our simulations, toroidal 360° grids are used and almost all relevant the plasma facing components, including those toroidal nonsymmetrical (especially the limiters), are taken into account. An example of the computational grids (poloidal cross-section) and the gas puffing positions for AUG and JET is shown in Fig. 1. Note that both in AUG and JET, the mid-plane gas source is essentially a poloidal and toroidal extended source filling the whole A-port. The divertor gas puffing is used as a reference case. Firstly the plasma of this reference case is validated with various experimental data by choosing the appropriate particle and energy transport coefficients. The separatrix density and total input power are set as the same as those in the experiments. Then the

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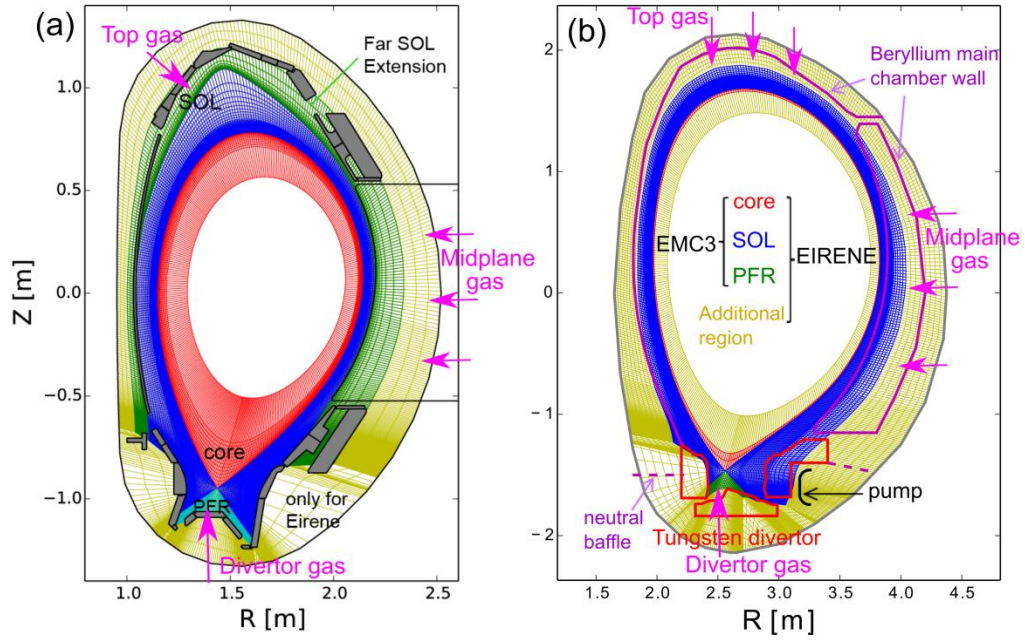


Fig. 1. Gas puffing positions and poloidal cross-section of the EMC3-EIRENE computational grid for (a) AUG; (b) JET. Figure reproduced with permission from [4, 5].

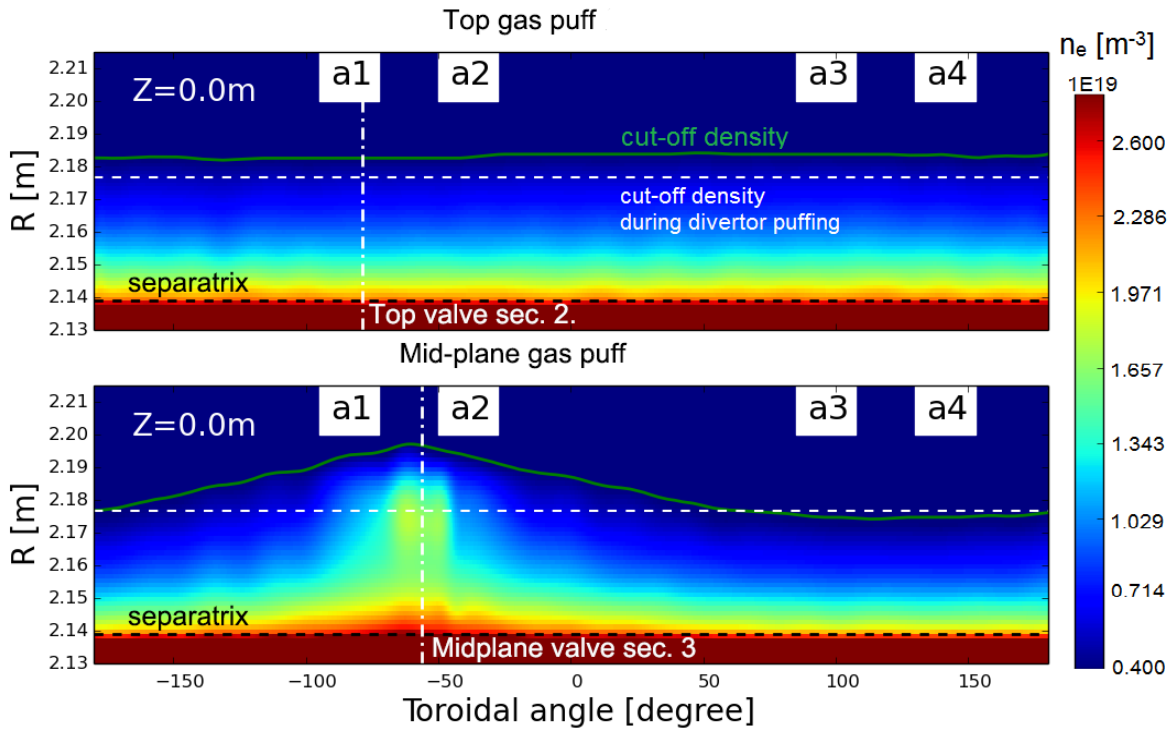


Fig. 2. Toroidal cross-sections (in the outer mid-plane) of the edge electron densities simulated by EMC3-EIRENE for top and mid-plane gas puffing in AUG H-mode plasmas. The white dashed line is the cut-off density during divertor gas puffing and is used as reference; the green line is the cut-off density during top or mid-plane gas puffing; the vertical dash-dotted lines represent the toroidal positions of the gas valves. The four ICRF antenna positions are represented by a1, a2, a3 and a4. Figure reproduced with permission from [4].

gas source is switched from the divertor to other locations of the machine to investigate other gas puffing cases.

Gas puffing from different locations of the machine ultimately lead to different SOL densities in front of the antenna. Such an example in AUG H-mode plasmas is shown in Fig. 2. Compared to divertor gas puffing, top gas puffing increases the SOL density near the outer mid-plane, where the antenna are located, almost

toroidally uniformly but to a small extent, and this density increase is independent of the toroidal position of the top gas valve. Mid-plane gas puffing increases the SOL density very significantly but locally, and this density increase depends sensitively on the toroidal position of the mid-plane gas valve. The largest density increase is found in locations close to the gas valve and this increase decays in the toroidal direction. Same phenomena are observed also in JET.

The mechanisms of the local density increase with local gas puffing are summarized as follows. Gas puffing provides a local particle source. In a steady state, the ionized particles (ions+electrons) leave this ionization zone mainly via convection along the field lines. The associated convective energy flux and the power losses into the ionization processes provide an energy sink to the background plasma and thereby lower the plasma temperatures in the ionization region due to the finite classical heat conductivity, resulting in a local density rise as a consequence of the parallel pressure conservation.

The 3D plasma density obtained from the EMC3-EIRENE simulations are then used in the antenna codes such as FELICE [11] and RAPLICASOL [9] to calculate the antenna resistance R_c . Good agreements are found between the calculated and measured values. For top gas puffing, the increase of R_c , $\sim 25\%$ in AUG and $\sim 40\%$ in JET H-mode plasmas, is almost the same for all the antennas. For mid-plane gas puffing, the largest increase of R_c , $\sim 120\%$ both in AUG [1] and JET [3] H-mode plasmas, is found for the antenna closest to gas valve. This value decays with the distance between the antenna and valve becomes larger: the decay length is $\sim 3.7\text{m}$ in AUG [1, 4] and $\sim 3.0\text{m}$ in JET [3]. The difference of the decay lengths in various machines can be attributed to the peculiarity of the wall structures (for instance JET has wide poloidal limiters), different magnetic field geometries and plasma transport properties. Nevertheless, the results in AUG and JET strongly suggest that mid-plane gas valves close to the antenna is the best option to maximize the ICRF power coupling.

3 ICRF induced plasma convection

The interaction between the ICRF waves and SOL plasma is a complex nonlinear process: (1) the plasma density influences the E_{\parallel} fields; (2) the E_{\parallel} fields lead to an inhomogenous Direct Current (DC) sheath potential, V_{DC} ; (3) the inhomogenous V_{DC} drives $E \times B$ drifts in the SOL and modify the plasma density. The new plasma density will change the E_{\parallel} fields, i.e. the loop comes to step (1) again.

A first simulation strategy is to use the measurements of the sheath potential of the reciprocating Retarding Field Analyzer (RFA) to estimate the $E \times B$ drifts, recently implemented in EMC3-EIRENE [8]. This makes possible to compare the impact of the $E \times B$ drifts on the density, predicted by EMC3-EIRENE, with the measurements of RFA or reflectometries. Our simulation results are in good agreement with the experimental ones, Fig. 3. The results indicate that convective cells are developed where high potential blobs exist. Density depletions are found in the center of the convective cells while density accumulations are found where the plasma interacts with the wall. The plasma flows driven to the wall by the convective drifts enhance the interaction between the plasma and wall.

The second simulation strategy is of the ab-initio type, and it requires to interface several codes in a consistency loop [12]. Before starting the iterative loop, EMC3-EIRENE simulations are done for a phase of the plasma discharge without RF, and thus without $E \times B$ drifts induced by ICRF. This is necessary to set the particle and transport coefficients of EMC3-EIRENE consistently with the experimental plasma profiles. These coefficients are assumed unaffected by ICRF heating and thus kept constant in the loop. When this setting is completed, the consistency loop for the ICRF heating phase starts by calling RAPLICASOL code to evaluate the RF E_{\parallel} fields. These fields are, then, used in SSWICH code to calculate V_{DC} , which, in turn, defines the $E \times B$ drifts for the call of EMC3-EIRENE. The iteration loop stops when the difference of a certain parameter (such as density) in one loop with that in the previous loop is lower than 5%. An example of this self-consistent simulation loop is shown in Fig.4. Three cases have been investigated: (a) 2-strap antenna with power balance between the left and right straps $P_{\text{left}}/P_{\text{right}}=1:1$ (the most typical settings for 2-strap antenna); (b) 3-strap antenna with power balance between the central and outer straps $P_{\text{center}}/P_{\text{outer}}=3:2$ (optimized power balance); (c) 3-strap antenna with $P_{\text{center}}/P_{\text{outer}}=1:9$ (non-optimized power balance). All the studied cases are with dipole phasing. The results (Fig. 4) indicate that the largest density convection happens in the top and bottom of the antennas. The largest sheath potential and density convection are induced by the 2-strap antenna (Case (a)) while the lowest values are induced by the 3-strap antenna with optimized power balance (Case (b)).

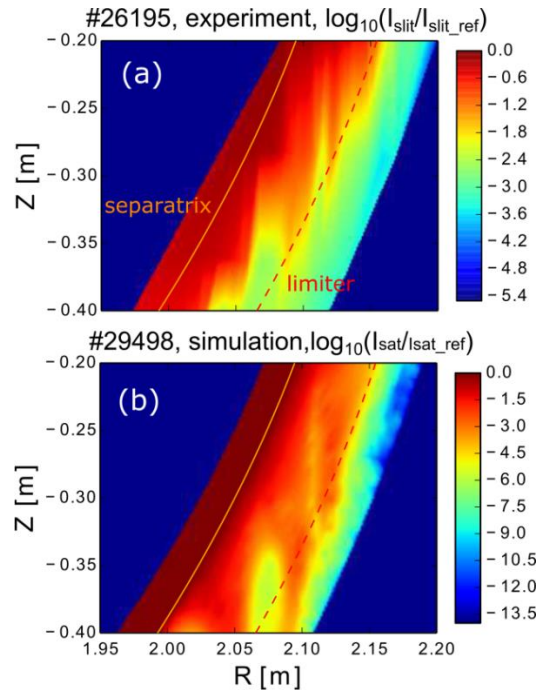


Fig. 3. Comparisons between experimental and simulated ion saturation current (in logarithmic scale). The ion saturation current can be expressed as $I_{\text{sat}}=qn_eC_s$, in which q is the ion charge number, n_e is the electron density and C_s is the plasma sound speed. Figure reproduced with permission from [8].

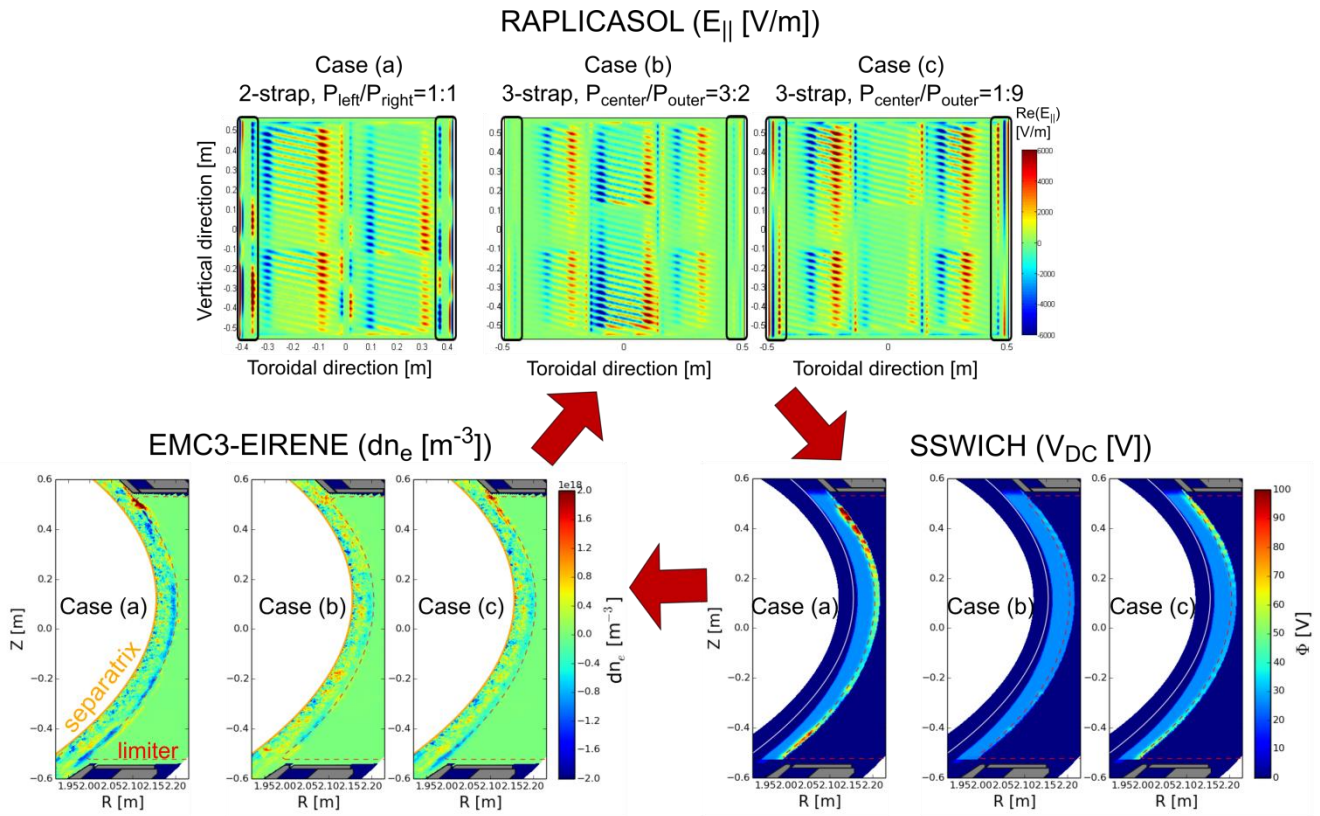


Fig. 4. Results of the self-consistent loop, including $E_{||}$ calculated with the RAPLICASOL code, V_{DC} calculated with the SSWICH code and n_e calculated with the EMC3-EIRENE code (here $dn_e = n_e - n_{e_ref}$, in which n_{e_ref} is the unperturbed density). Figure reproduced with permission from [12].

4 Conclusions

Through edge plasma modelling with ICRF coupling, significant progresses have been made on understanding (1) the effects of local gas puffing on edge plasma density and ICRF coupling; (2) the ICRF induced density convection.

The 3D SOL plasma density during different gas puffing are simulated with the EMC3-EIRENE code, which are then used in the antenna codes to calculate the coupling resistances. Different gas puffing cases in AUG and JET have been extensively studied and extensive comparisons between the simulations and experimental results have been made. It is shown that the gas valves close to the antennas increase the local density and thus the ICRF power coupling most significantly. This conclusion is expected to be also true for other and future machines. Similar numerical studies for ITER and DEMO are being carried out to figure out optimized gas valve positions for the best ICRF power coupling.

The ICRF induced density convection has been studied either by EMC3-EIRENE simulations based on experimental data or by self-consistent simulations with several codes (EMC3-EIRENE + RAPLICASOL + SSWICH). Our simulation results are in qualitative agreement with the measurements. It is indicated that convective cells are often developed in the top and bottom of the antennas. The convective drifts change the density in front of the antenna locally, influence the

ICRF power coupling and enhance the plasma-wall interactions. Among the studied cases, the 3-strap antenna with optimized feeding configurations leads to the lowest rectified sheath potential and density convection. The developed numerical methods can be further used to predict and optimize the experiments.

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