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## De-Convolution of Complex Residual Gas Spectra at JET

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\* See annex of F. Romanelli et al, "Overview of JET Results",

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#### ABSTRACT.

Spectral analysis of the vacuum conditions within the JET tokamak provides unique challenges due to the complex interactions of Hydrogen, Deuterium and Tritium with the large amounts of Carbon forming the plasma facing first wall. Overlapping masses and low resolution of the quadrupole Residual Gas Analyzer (qRGA) means discrimination of individual species is not possible operating in a conventional mode. Baking the vessel to temperatures in excess of 300 degrees Celsius further complicates an already difficult task.

Here, we present and demonstrate a complimentary technique operating the qRGA in a mode where the energy of the electrons emitted within the ionization source is variable. Different elements have defined ionization energies required to remove an orbiting electron, this energy is dependent on the electron orbital, i.e. outer shell electrons generally have weaker ionization energies due to the greater distance and lower electrostatic forces from the nucleus. Using this technique known as "Soft Ionization" de-convolution of the mass spectra obtained at JET could lead the way to providing a better understanding of the chemistry within the vacuum vessel as well as providing invaluable diagnostic information during vessel conditioning.

Initial experimental data has proved encouraging demonstrating ionization energy discrimination of  $D_2$  and  $^4$ He at 4amu and  $D_2$ O/Ne/Ar $^{++}$  at 20amu. De-convolution of hydrocarbon and deuterated hydrocarbon molecules in the mass spectrum is also currently ongoing, with initial results revealing clear discrimination of  $CH_4$  and  $CD_2$  molecular ions at 16amu. Using electron ionization cross-section theory we also demonstrate the generation of algorithms within the qRGA software to automatically discriminate quantitatively between overlapping peaks in the mass spectra. This technique can be applied to historical as well as real time data.

2010 brings the installation of the Beryllium plasma facing wall at JET and an opportunity to accurately determine the vessel chemistry during commissioning. The results from this phase of JET operations and further applications of the technique including leak detection and cryogenic pumping regeneration inventory determination could prove valuable for ITER commissioning and operations.

#### 1. INTRODUCTION

At JET, fusion fuel most commonly used in the tokamak is deuterium( $D_2$ ). Purity of the fuel is vital to ensure the data obtained from plasma pulses is useful, and that pulse recipe repeatability is not compromised.

Our primary goal was to determine the purity of the fuelling gas, deuterium, and confirm that we did not have any contamination from helium (He). Analysis using our quadrupole Residual Gas Analyser (qRGA) in the conventional way was not possible due to the overlapping masses of the two species at 4 atomic mass units (amu).

Here, we present and demonstrate the approach we used to solve this problem; a complimentary technique, Threshold Ionization Mass Spectrometry (TIMS), operating the qRGA mass spectrometer

in a mode allowing control over the energy of the electrons emitted within the ionization source. Recognising the potential to identify more complex compounds which have brought uncertainty when attempting to interpret data from the torus primary vacuum vessel we have applied this technique across a broad range of gases that exhibit overlapping masses in the laboratory. The JET residual signature is unique in that the combination of three isotopes of hydrogen: hydrogen, deuterium and tritium forming three isotopes of water: H<sub>2</sub>O, heavy water, D<sup>2</sup>O and super heavy water, T2O in addition to the 6 tonnes of carbon tiles forming the plasma facing wall resulting in complex carbon / hydrogen isotope interactions.

Baking the torus vacuum vessel to  $320^{\circ}$ C further complicates and already difficult task due to disorption of previously formed  $C_xH_v$  and  $C_xD_v$  species from the vessel wall.

#### 2. EXPERIMENTAL DETAILS OF TIMS AT JET

#### 2.1 OVERVIEW

The Hiden Analytical qRGA system at JET is operated in a mode where the energy of the electrons emitted within the ionization source is variable. Different elements have defined ionization energies required to remove an orbiting electron. This energy is dependent on the electron orbital, i.e. outer shell electrons generally have weaker ionization energies due to the greater distance and lower electrostatic forces from the nucleus. This gives rise to the electron impact "threshold ionization energy" curve. The ionization process of neutral particles commences at a minimum (threshold) energy of the impacting electrons.

This minimum energy is dependent and unique to any species present in the gas matrix, resulting in a spectral "identifier" or fingerprint for all atomic or molecular species.

### 2.2. EXPERIMENTAL METHOD

The Hiden Analytical HAL201 RC qRGA mass spectrometer was interfaced to a small test analysis chamber test chamber pumped via a turbomolecular pump at JET. A custom gas inlet system allowing the introduction into the analysis chamber of up to 10 different gases controlled by a calibrated mass flow controller (mfc) allows accuracy to < 1%. During this experiment the gas mix was ultra pure (100%)  $D_2$ , (100%)  $^4$ He, and small (1%) concentrations of  $CH_4$ ,  $C_2H_4$ ,  $C_2H_6$ ,  $C_3H_6$ ,  $C_3H_8$ ,  $C_4H_{10}$  in a (94%)  $N_2$  carrier gas.

The total system pressure of the analysis chamber was  $\sim 10^{-7}$  mbar during data collection.

## 2.3. EXPERIMENTAL AIMS

A gas mixture of  $D_2$  and <sup>4</sup>He was injected into the test chamber at JET with control of the ratio of the mixture with MFC's. We investigate the possibility of de-convolution of  $D_2$  from <sup>4</sup>He at 4amu in the mass spectrum, and real-time quantification analysis techniques using the qRGA TIMS technique, combined with BEB (Binary-Encounter-Bethe) ionization cross section (s) theory.

Additionally, a mix of  $D_2$  with the hydro-carbon mixture investigated the possibility of TIMS

de-convolution of complex  $H_xD_y$ ,  $H_xC_y$ ,  $C_xD_y$ , and preliminary real time qualitative data analysis. Finally we looked for evidence of  $D_2O$ , argon  $(Ar^{++})$  and neon (Ne) formation and de-convolution at 20amu.

## 3. RESULTS: TIMS DE-CONVOLUTION OF GAS MIXTURES AT JET

## 3.1. HELIUM <sup>4</sup>HE AND DEUTERIUM D<sub>2</sub> AT 4 AMU

Figures 1(a,b) shows real time raw qRGA TIMS data [at 4amu] of a <sup>4</sup>He/D<sub>2</sub> gas mixture displayed within qRGA software.

Figure 2 shows real time mass flow controlled ratio of  ${}^{4}\text{He/D}_{2}$  gas mixture displayed within qRGA software. The smaller inset in fig.2. shows the corresponding raw qRGA TIMS data [at 4amu] taken as a function of time.

The two traces, in the inset of fig.2. represent  $N_{(D_2)}$  and  $N_{(D_2 + He)}$  at 19eV and 31eV respectively. The real time qRGA data on the right shows the true  ${}^4\text{He/D}_2$  abundance ratio  $n_{(D_2)}/n_{(He)}$  after applying BEB theory automatically calculated and displayed in the qRGA software.

Application of the BEB threshold ionization cross section theory to qRGA data: From theory,  $N_{(D_2)}$  is the  $D_2$  partial pressure signal measured at 19eV and  $N_{(D_2 + He)}$  is the  $D_2$  and  $^4$ He partial pressure signal combined, measured at 31eV in the qRGA. Now,  $N_{(D_2)}$  and  $n_{(He)}$  are the real abundance densities of  $D_2$  and  $^4$ He in the sample gas. To obtain the ratio,  $n_{(D_2)}/n_{(He)}$  from the measured partial pressure signals  $N_{(D_2)}$  and  $N_{(D_2 + He)}$ , to a first approximation it can be shown;

$$N_{(D_2)} = A_i \left( \sigma_{(D_2)}^{\{19eV\}} \right) \cdot \eta_{(D_2)} \tag{1}$$

and similarly,

$$N_{(D_2 + {}^4He)} = A_i \left\{ \left( \sigma_{(D_2)}^{\{31eV\}} \right) \cdot \eta_{(D_2)} + \left( \sigma_{({}^4He)}^{\{31eV\}} \right) \cdot \eta_{({}^4He)} \right\}$$
 (2)

Ai is the instrument (qRGA) factor which is the same for both  $D_2$  and  $^4$ He as the signal is measured at 4amu. It is also independent of electron ionization energy between energies of 19 and 31eV. Also, we define the ionization cross sectional coefficients for  $D_2$  and He as  $N_{(D_2)}$  and  $n_{(He)}$  respectively: Solving (1) and (2) yields;

$$\frac{\eta_{(D_2)}}{\eta_{(^4He)}} = \frac{\left(\sigma_{(^4He)}^{\{31eV\}}\right)}{\frac{N_{(D_2+^4He)}}{N_{(D_2)}} \left(\sigma_{(D_2)}^{\{19eV\}}\right) - \left(\sigma_{(D_2)}^{\{31eV\}}\right)}$$
(3)

### 3.2. HEAVY WATER, D2O AND DOUBLY CHARGED ARGON AT 20 AMU.

Figure 3. shows separation of small (<1%) quantities of heavy water from chamber residual Argon. It can also be seen in Fig 4. that TIMS clearly identifies formation of HD+ molecule at 3amu via the onset at 15.5eV.

This is an important result as it will allow confident de-convolution of HD from T+ (3amu) when Tritium is introduced into the JET fuel mixture during future Tritium campaigns at JET. (note the inset shows the derivative of the TIMS HD+ curve to qualitatively further enhance species deconvolution in the TIMS spectra).

#### 3.3. HYDROGEN AND DEUTERIUM INTERACTIONS WITH CARBON AT 16 AMU

Preliminary results are shown in Fig.5 of interaction of Deuterium ( $D_2$ ) and hydrocarbons resulting in the formation of  $C_xH_y$  and  $C_xD_y$  molecules. For example at 16amu, overlapping species are  $CH_4^+$  and  $CD_2^+$ . TIMS onsets are 12.3eV and 18.24eV respectively. Below species de-convolution is shown with TIMS. Again the differential;

$$\frac{dy}{dx} = \frac{d(partial pressure)}{d(electron energy)} ,$$

is shown simply to qualitatively enhance the de-convolution of the transition regions, with the minima in the corresponding to the transition from  $CH_4^+$  to  $(CH_4^+ + CD_2^+)$ .

In figure 5, the hydro-carbons in the  $50:50 D_2/(N_2 \text{ carrier gas})$  are present at just a 0.5% fraction of the gas mixture total composition, which demonstrates TIMS viability as a use for low level impurity detection.

## 3.4. HEAVY WATER, D<sub>2</sub>O AND NEON (NE) AT 20 AMU.

Neon (Ne) at mass 20 has also been separated in this way. The two spectra above, Fig.6, shows the raw data curve (top) and the derivative (bottom) clearly demonstrating the separation of the  $D_2O$  at 12.6eV and Ne at 21.5eV, both sharing the same mass at 20amu. Separation of these masses will prove a useful tool at JET as the vessel interspaces are filled with neon so detection of a leak from one of these into the torus can be identified easily.

## CONCLUSIONS AND FUTURE WORK

Whilst here we have demonstrated findings from some of our first laboratory experiments our success has identified characteristics of many of the species in Table 1. Occasionally overlapping species may have first ionization energies too close to each other to distinguish with our equipment. These cases are easily overcome but utilising the 2nd or 3rd ionization energies by increasing the energy of the electrons emitted within the qRGA source to eject orbiting electrons from more tightly bound shells.

Using the qRGA TIMS technique, de-convolution of the mass spectra obtained at JET could lead the way to providing a better understanding of the chemistry within the vacuum vessel, as well as providing invaluable diagnostic information during vessel conditioning.

Initial TIMS has proved encouraging, demonstrating discrimination of  $D_2/^4$ He (4amu) and  $D_2O/Ar^{++}$  (20amu). De-convolution of  $C_xH_y$  and  $C_xD_y$  molecules in the mass spectrum is ongoing, with initial resultsde-convoluting  $CH_4^+$  and  $CD_2^+$  (16amu). With application of BEB to TIMS, we have

demonstrated the generation of algorithms within the qRGA software to automatically discriminate different species quantitatively with overlapping peaks in the mass spectra. Further experimental work will continue using our dedicated gas inlet with the addition of a carbon reaction cell to study the dissociation of species with temperature as well as generate  $C_xH_y$  and  $C_xD_y$  species to obtain further TIMS data for the purpose of improving the technique.

In addition to this the installation of the Hiden HAL201 RC qRGA is taking place on the JET Gas Introduction System Matrix and on the JET torus. From these instruments we hope to prove our de-convolution algorithms to give an accurate breakdown of the composition of the complex residual gas spectra at JET.

#### **ACKNOWLEDGMENT**

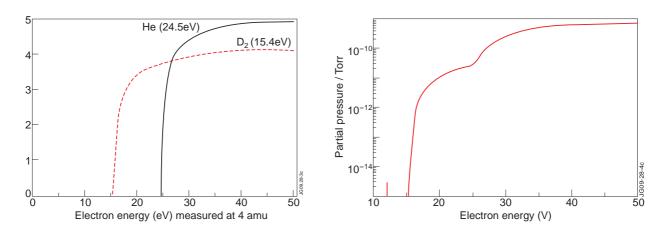
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- [2]. Yong. K, J. Res. Nat. Inst. Stand. Technol. **105**, 285. (2000)

| Species | Major m/z | %   | Major1 m/z | %  | Major2 m/z | %  | Major3 m/z | %  | Major4 m/z | %  | Major5 m/z | %  | Major6 m/z | %  | Rel Sens | 1st IE (eV)  |
|---------|-----------|-----|------------|----|------------|----|------------|----|------------|----|------------|----|------------|----|----------|--------------|
| H2      | 2         | 100 |            |    |            |    |            |    |            |    |            |    |            |    |          | 15.42        |
| He      | 4         | 100 |            |    |            |    |            |    |            |    |            |    |            |    | 0.14     | 24.58        |
| D2      | 4         | 400 |            |    |            |    |            |    |            |    |            |    |            |    |          | 15.46        |
| CH4     | 16        | 100 | 15         | 85 | 14         | 16 | 13         | 5  | 12         | 2  | 17         | 1  |            |    | 1.6      | 12.61        |
| H2O     | 18        | 100 | 17         | 21 | 16         | 2  |            |    |            |    |            |    |            |    | 0.9      | 12.61        |
| D2O     | 20        | 100 | 18         | 21 | 16         | 2  |            |    |            |    |            |    |            |    | 0.9      |              |
| CD4     | 20        | 100 | 18         | 85 | 16         | 16 | 15         | 5  | 14         | 2  |            |    |            |    | 1.6      | 12.65        |
| Ne      | 20        | 100 | 22         | 10 |            |    |            |    |            |    |            |    |            |    |          | 21.56        |
| Air     | 28        | 100 | 32         | 27 | 14         | 6  | 40         | 1  |            |    |            |    |            |    | 1        |              |
| CO      | 28        | 100 | 12         | 5  | 16         | 2  |            |    |            |    |            |    |            |    | 1.05     | 14.01        |
| C2H6    | 28        | 100 | 27         | 33 | 39         | 26 | 26         | 23 | 29         | 21 | 15         | 5  |            |    | 2.6      | 11.57        |
| C2D2    | 28        | 100 | 26         | 17 | 14         | 5  | 24         | 4  | 12         | 1  |            |    |            |    |          | 11.2         |
| C3H8    | 28        | 100 | 26         | 59 | 27         | 41 | 44         | 28 | 43         | 23 | 39         | 19 | 41         | 13 |          | 10.94        |
| C2H4    | 28        | 100 | 27         | 64 | 26         | 53 | 25         | 8  | 14         | 2  |            |    |            |    |          | 10.51        |
| N2      | 28        | 100 | 14         | 15 |            |    |            |    |            |    |            |    |            |    |          | 15.58        |
| C2D4    | 32        | 100 | 28         | 64 | 30         | 62 | 16         | 11 | 26         | 10 | 14         | 6  | 12         | 4  |          | 10.54        |
| C2D6    | 32        | 100 | 30         | 24 | 36         | 20 | 28         | 18 | 34         | 16 | 18         | 4  | 16         | 4  |          | 11.73        |
| 02      | 32        | 100 | 16         | 22 |            |    |            |    |            |    |            |    |            |    |          | 12.06        |
| C3D8    | 34        | 100 | 32         | 60 | 30         | 30 | 50         | 22 | 42         | 16 | 46         | 12 | 28         | 7  |          | 11.4         |
| Ar      | 40        | 100 | 20         | 16 |            |    |            |    |            |    |            |    |            |    | 1.2      | 15.75        |
| C3H4    | 40        | 100 | 39         | 80 | 38         | 30 | 37         | 20 | 36         | 5  | 41         | 4  |            |    |          | 10.36        |
| C3H6    | 42        | 100 | 39         | 73 | 44         | 70 | 27         | 39 | 40         | 29 | 38         | 20 | 37         | 13 |          | 9.73         |
| C4H10   | 43        | 100 | 29         | 44 | 27         | 40 | 28         | 32 | 41         | 40 | 39         | 14 | 58         | 13 |          | 10.53        |
| CO2     | 44        | 100 | 16         | 9  | 28         | 8  | 14         | 8  | 22         | 3  |            |    |            |    | 1.4      | 13.79        |
| C3D4    | 44        | 100 | 16         | 89 | 40         | 35 | 38         | 27 | 36         | 8  | 26         | 4  |            |    |          | 10.37        |
| C3D6    | 46        | 100 | 42         | 74 | 48         | 70 | 30         | 39 | 44         | 30 | 40         | 20 | 38         | 13 |          | 9.76<br>10.5 |
| C4D10   | 50        | 100 | 34         | 44 | 30         | 37 | 32         | 37 | 46         | 28 | 42         | 13 | 48         | 12 |          | 10.5         |

Table 1:



Figures 1(a,b): Real time raw qRGA TIMS data [at 4amu] of a  $^4$ He/D<sub>2</sub> gas mixture displayed within qRGA software.

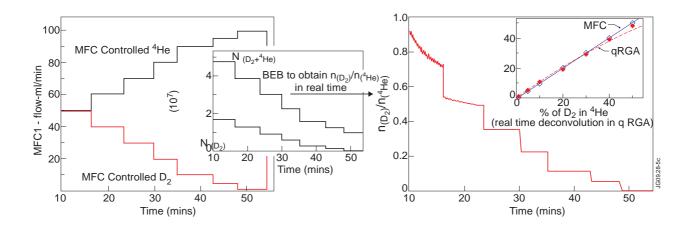
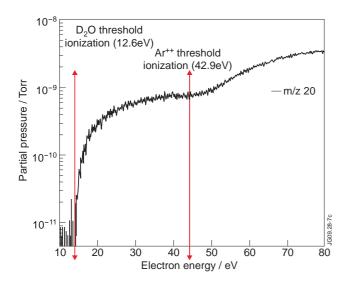


Figure 2: Real time mass flow controlled ratio of  $^4$ He/D $_2$  gas mixture displayed within qRGA software. The smaller inset in fig.2. shows the corresponding raw qRGA TIMS data [at 4amu] taken as a function of time.



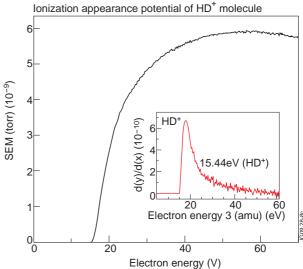


Figure 3: Separation of small (<1%) quantities of heavy water from chamber residual Argon.

Figure 4: TIMS showing formation of HD+ molecule at 3amu via the onset at 15.5eV.

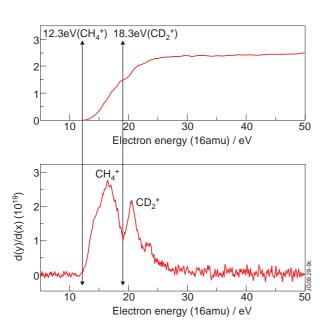


Figure 5: Preliminary results of interaction of Deuterium  $(D_2)$  and hydrocarbons resulting in the formation of  $C_xH_y$  and  $C_xD_y$  molecules.

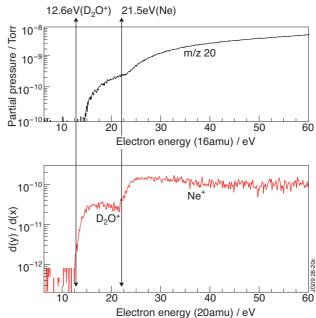


Figure 6: Raw data curve (top) and the derivative (bottom) demonstrating the separation of the  $D_2O$  at 12.6eV and Ne at 21.5eV, both sharing the same mass at 20amu.